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The President's Report

July 1, 2011 - June 30, 2012

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"... to encourage, in the broadest and most liberal manner, investigation, research, and discovery, and the application of knowledge to the improvement of mankind"

The Carnegie Institution was incorporated with these words in 1902 by its founder, Andrew Carnegie. Since then, the institution has remained true to its mission. At six research departments across the country, the scientific staff and a constantly changing roster of students, postdoctoral fellows, and visiting investigators tackle fundamental questions on the frontiers of biology, earth sciences, and astronomy.

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The President's Commentary





Carnegie president Richard A. Meserve Image courtesy Jim Johnson

¹See, e.g., President's Council of Advisors on Science and Technology, *Report to the President, Transformation and Opportunity: The Future of the U.S. Research Enterprise*, 3 (Washington, D.C.: Executive Office of the President, 2012), pp. 19-21; Committee on Science, Engineering, and Public Policy, *Rising Above the Gathering Storm: Energizing and Employing America for a Brighter Future* (Washington, D.C.: National Academies Press, 2007), pp. 41-67. e live in turbulent economic times. As I compose this text, Congress has temporarily avoided the fiscal cliff that loomed in early January but has signaled that further and even more difficult conflicts must be resolved in the coming months. Against this backdrop, we should be worried about the future trajectory of our economy. But even if these problems are successfully resolved, there is one troubling certainty: significant downward adjustments in federal funding for the discretionary part of the federal budget is inevitable for many years. Even with growing tax revenues, there is simply no way that our country's debt obligations can be met and that our entitlement programs, even with adjustments, can be funded without significant budget reductions elsewhere. Although both political parties support scientific research, powerful advocates support other parts of the federal budget. We must worry about how support for science will fare in the inevitable radical restructuring of federal spending.

A strong case can be made for the sustenance of basic science even in these troubling economic times. Numerous studies over the years have consistently shown that investments in scientific research pay off in significant ways. The ground-breaking work of Robert Solow—work that garnered him a Nobel Prize—showed that more than half of the productivity growth in the first half of the 20th century could be attributed to technological change. Subsequent studies by others have confirmed the fundamental importance of research and development (R&D) as an engine for economic advance.¹

The fact that R&D provides a healthy return on investment does not, of course, by itself justify federal investment. A significant share of total U.S. R&D expenditures is (and should be) undertaken by the business sector—about 71% in 2009.² But the business sector invests far less in basic research than is socially optimal because the returns from such research are very uncertain and cannot necessarily be captured by the entity that makes the investment. Indeed, the pressure for short-term returns has meant that business-sector investment in basic research has declined over recent decades, and once vital industrial research centers, such as Bell Labs and RCA Labs, no longer focus on basic research. Industry predominantly directs its support to applied or developmental activities, which promise near-term returns that can be captured by the funder.

In light of these facts, the government has a critical role in the support of basic research for the benefit of all. In 1945 Carnegie President Vannevar Bush made exactly this argument to President Truman in the report *Science—The Endless Frontier*. This report ultimately resulted in the establishment of the National Science Foundation, launched the dramatic rise in federal support for basic

²National Science Board, *Science and Engineering Indicators 2012* (Washington, D.C.: National Science Foundation, 2012), Table 4-1.

research over the past 68 years, and is largely responsible for the university research enterprise that has proven to be an important source of new discoveries.

Of course, by its very nature, the returns from any specific project in basic research are difficult to predict. The whole point of basic research is the exploration of the unknown, driven by an individual scientist's curiosity and quest for knowledge. What is remarkable is that this type of "unfocused" work can open entirely new vistas for advance that could not be contemplated previously and of which even the scientist was unaware. Basic research in quantum mechanics provided the understanding on which today's microelectronics, communications, and computer businesses were built. Basic work in mathematics and computer science laid the foundations for the Internet, spawning some of the world's largest companies. And even less earth-shattering work can provide for the steady advance of products and services. Because it is sometimes impossible to determine before the fact what research will contribute in this way, the aim should be to support a wide portfolio in anticipation that astonishing and often unanticipated gains will result. For this reason alone, significant federal support for basic research should be maintained. We will regret the consequences if we starve the goose that lays the golden eggs.





Of course, many of the benefits derived from basic science extend beyond the economic effects. Basic science is the underpinning for advances in healthcare, contributing to both the length and quality of life. It underlies our national security, providing deterrence and a capacity to respond to both conventional warfare and the difficult threats of terrorism, cyberwarfare, and biological attacks. It provides the capacity to respond to climate change and to the challenges of providing adequate energy supply, feeding the growing world's population, achieving sustainable economies, protecting the environment, and meeting the emerging challenge of providing sufficient potable water. Perhaps, most fundamentally, basic science satisfies a deep-seated human desire to *know*—to understand the universe and our place in it.

As revealed in the subsequent pages of this Year Book, the work of Carnegie scientists exemplifies the promise of basic research. Consider the following examples:

Steve Farber of our Department of Embryology is undertaking path-breaking work on the metabolism of fats and cholesterol. Using fluorescently tagged lipids, he is able to observe absorption in the small intestine and subsequent metabolism in live zebrafish, a remarkable model animal for such work because





Ronald Cohen (above) and former intern Maimon Rose discovered a new and efficient way to pump heat at miniscule scales using crystals. The image at left shows a molecular dynamics simulation of lithium niobate under a time-varying electric field. The electric field changes the sign of the polarization—the critical component to the discovery. Niobium is red, oxygen is green, and lithium shows a range of colors for different time steps. The niobium and oxygen are shown for one time step only, for clarity. zebrafish are transparent in their early stages of development. The work has important implications for our understanding of a variety of ailments, including diabetes, obesity, and cardiovascular disease.

Ron Cohen of our Geophysical Laboratory and Maimon Rose, a former highschool Carnegie intern, have conducted work on ferroelectric crystals—materials that have electrical polarization in the absence of an electric field. Applying an external electric field reverses the polarization and causes a temperature change in the material. They found a very dramatic temperature change in their studies of ferroelectric lithium niobate. The work holds the prospect that such crystals could be used to pump or extract heat. The crystals might find application on computer chips to prevent overheating, which currently limits higher computing speeds.

David Ehrhardt of the Department of Plant Biology and Ryan Gutierrez have been studying cellulose, the crucial component of plant cell walls. Normally the individual chains that make up cellulose bond to each other to make a semi-crystalline fiber, which provides a plant with rigidity and strength. This fiber is also responsible for cellulose's resistance to digestion. In collaboration with Seth DeBolt



Diana Roman studies how magma travels through the Earth's crust and in volcanic conduit systems to understand seismicity and stress changes as magma moves. She is deploying a broadband seismometer at the Crater Lake volcano in Oregon. at the University of Kentucky, a former Carnegie postdoctoral fellow, the researchers have discovered mutations in the genes that encode the cellulose-making proteins, resulting in plants with cell walls with structural defects. This results in the production of cellulose that is less recalcitrant to digestion. The work may provide a pathway to liberate sugars from cellulose, a crucial step in the production of biofuels.

Diana Roman of the Department of Terrestrial Magnetism is studying the formation, structure, evolution, and dynamics of the conduit systems for the transport of magma in volcanoes, and the relationship of these systems to the microearthquake swarms that occur in the vicinity of active volcanoes. The work will allow greater understanding of volcanoes, with the promise of eventually enabling prediction of the timing and scale of eruptions.

Global Ecology's Greg Asner and his team are using the Carnegie Airborne Observatory (CAO) to revolutionize wide-scale ecological studies. The researchers combine laser and spectral imaging technology onboard a twin-engine aircraft to derive simultaneous measurements of an ecosystem's chemistry, structure, biomass, and biodiversity. In just one year, the system has mapped tens of millions of acres



Greg Asner's team recently rolled out results from the new Airborne Taxonomic Mapping System (AToMS) mounted on the Carnegie Airborne Observatory (CAO), a fixed-wing aircraft (top right). The groundbreaking technology and its scientific observations are uncovering a previously invisible ecological world. The CAO landscape shows the height of vegetation, with red areas tallest.

of ecosystems in California, Panama, Colombia, Costa Rica, and the Peruvian Amazon basin. In addition to enhancing the understanding of ecosystems, the technology has applications for mitigating climate change, forest conservation, and ecosystem management. Indeed, the CAO could be a crucial element in the implementation of the United Nations initiative Reducing Emissions from Deforestation and forest Degradation (REDD) by enabling the swift and accurate measurement of carbon stocks in protected forests.

A team of astronomers lead by **Wendy Freedman, the director of the Observatories**, has used NASA's Spitzer Space Telescope to make one of the most accurate measurements of the Hubble constant yet achieved. The Hubble constant is a measure of the universe's expansion rate, and it underlies our understanding of the universe's age and size. By establishing a precise measure of the rate at which the universe is expanding at the current time, the research provides insights into the observed acceleration of the expansion rate over time—perhaps the most fundamental scientific mystery of our age.

Finally, as an adjunct to our research activities, our educational activities reflect the same innovation we bring to our scientific research. The staff of the Carnegie Academy for Science Education (Julie Edmonds, Toby Horn, and Marlena Jones)





has worked closely with Friendship Collegiate Academy to develop a new teaching approach that integrates instruction in science, technology, engineering, and math with English and social studies. The Washington, D.C., branch of Math for America, which was launched by Carnegie and American University—with careful shepherding by my predecessor Maxine Singer, continues to develop master teachers in mathematics for Washington, D.C., schools.

As these examples reveal, Carnegie remains an oasis of exciting activities in an uncertain time. Indeed, although the threat of reduced federal support looms for all of science, Carnegie has fared through the recent turmoil rather successfully. Our federal grants and contracts have been sustained in a period when many have endured cutbacks-a tribute to the capability of the Carnegie scientists. And unlike many institutions and universities, we have the benefit of balanced support for research from additional sources-our endowment, gifts from foundations and individuals, and other income (e.g., from rentals of our P Street headquarters). Thanks to the careful stewardship by our Finance Committee, our endowment had a gain of 5.4% over the 2011-12 fiscal year, which is well above the State Street median for endowments and foundations for the period. (Indeed, the Carnegie endowment has outpaced the State Street median over the two-, three-, and five-year periods as well.) And our efforts to strengthen our outreach to the broader community have served many purposes, including a growing flow of support for Carnegie science. At the same time, our frugality and care in operations has gained us the highest rating available from Charity Navigator, the nation's largest evaluator of non-profits and charities for fiscal management; we are one of only five non-profit organizations of the 5,500 that are monitored to have achieved this ranking \$13.1 Million for twelve years running. In short, although I observe the overall financial climate with concern, Carnegie has shown that it has the capacity to thrive. With your continuing help and support, we will do so.

Richard A. Meserve, President



Friends, Honors & Transitions

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Carnegie Friends

Annual Giving

Gifts were received between July 1, 2011, and June 30, 2012.

The Barbara McClintock Society

An icon of Carnegie science, Barbara McClintock was a Carnegie plant biologist from 1943 until her retirement. She was a giant in the field of maize genetics and received the 1983 Nobel Prize in Physiology/Medicine for her work on patterns of genetic inheritance. She was the first woman to win an unshared Nobel Prize in this category. To sustain researchers like McClintock, annual contributions to the Carnegie Institution are essential. The McClintock Society thus recognizes generous individuals who contribute \$10,000 or more in a fiscal year, making it possible to pursue the highly original research for which Carnegie is known.



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Carnegie Institution for Science



★ Robert and Margaret Hazen

Robert and Margaret Hazen

ver the years, Geophysical Laboratory (GL) scientist Bob Hazen and his wife, Margee, have become among the most ardent supporters of Carnegie science. Bob came to Carnegie as a postdoctoral fellow in 1976 and has been a staff researcher since 1978. His work is expansive—from the origins of life and the emergence of prebiotic chemistry to mineral evolution. In addition to his Carnegie research, Bob is the Clarence Robinson Professor of Earth Science at George Mason University. He lectures around the world and is dedicated to public education and outreach on numerous scientific topics. Bob has authored more than 350 scientific articles and 20 books on science, history, and music, many in collaboration with Margee.

The Hazens' contributions to Carnegie go well beyond Bob's extraordinary science. In 1989 they began donating annually to support scientific research and instrumentation at GL. They started an instrument fund in honor of the late GL scientist Thomas Hoering. In 2001 the Hazens made significant contributions to the Carnegie Campaign for Science. Since then they have made annual donations to support GL fellowships, interdisciplinary studies, renovations to the Broad Branch Road campus, and more. They are members of the Edwin Hubble and Second Century Legacy Societies.

Margee's support for Carnegie science also exceeds the norm. She was pivotal in the institution's centennial celebrations in 2002. An accomplished author and historian, she curated the centennial exhibition *Our Expanding Universe* and coauthored the book *Good Seeing*; both highlighted the first century of Carnegie's remarkable scientific discoveries, and neither would have succeeded without her long hours and unwavering dedication.

More recently, Bob has been forging a new, interdisciplinary field involving researchers worldwide: the study of the Earth's deep carbon cycle. It began in 2007 when Bob gave a lecture in New York about understanding life's geochemical origins and attracted the attention of Jesse Ausubel, a science advisor at the Alfred P. Sloan Foundation. Today Bob serves as executive director of the Deep Carbon Observatory, which has important Sloan funding, among his numerous other duties and research projects. Carnegie is extraordinarily fortunate that Bob and Margee have chosen to support the institution in so many different ways. We are deeply indebted to them for their enduring commitment.

Burton and Deedee McMurtry

ith an electrical engineering background and a passion for astronomy, Burton McMurtry was a natural fit for Carnegie. Trustee William Hearst introduced Burton to the institution, by way of their mutual interest in astronomy. Burton's first-hand exposure to Carnegie came through a trip that he took with his wife, Deedee, to the Las Campanas Observatory in early 1995. He was captivated by the experience and was elected to Carnegie's board in December 1996.

Burton received two bachelor's degrees from Rice University and master's and doctoral degrees in electrical engineering from Stanford University. He then worked for GTE Sylvania in engineering and management until 1969, when he changed course and went into the venture capital business. Burton founded several venture firms, including Technology Venture Investors (TVI), with a focus on start-up electronic companies. Among the many successful companies that his firms backed were ROLM Corporation, KLA-Tencor, Intuit, Microsoft, Sun Microsystems, and VeriFone. In 1995 Burton began retiring from TVI, and in 2004 he became the chairman of the Stanford University board.

Initially, Burton served on Carnegie's Research committee. In 1999 he became a member of the Finance committee, where he provided invaluable guidance on Carnegie's investments.

Burton generously supported Carnegie even before he joined its board. His first gift, in 1996, was for the wide-field camera for the Magellan Telescope Project. Later Burton served on the Visiting committee for the Observatories. His interests expanded to include Global Ecology, where he served on the committee that selected the architectural firm for the department's innovative "green" building. Burton and Deedee have generously donated every year to the Annual Fund. They have additionally supported the Carnegie Campaign for Science, Science for the City, and the Department of Embryology's Singer Building. They are members of the Hubble Society.

Among his many other contributions, Burton initiated a flourishing relationship with the Gordon and Betty Moore Foundation. In 2001 he met with Gordon Moore and discussed the then-new Department of Global Ecology. Since 2002, the Moore Foundation has made some \$13 million in grants to support work at Global Ecology. Carnegie is deeply grateful for Burton McMurtry's leadership and guidance over the years and for the couple's generous, steadfast support of many facets of Carnegie science.



★ Burton and Deedee McMurtry

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Foundations and Corporations

\$1,000,000 or more

The John D. and Catherine T. MacArthur Foundation The Cynthia and George Mitchell Foundation Gordon and Betty Moore Foundation Alfred P. Sloan Foundation

\$100,000 to \$999,999

The Ahmanson Foundation Blue Moon Fund, Inc. The Margaret A. Cargill Foundation The Gayden Family Foundation Michael E. Gellert Trust The Grantham Foundation for the Protection of the Environment Richard Lounsbery Foundation, Inc. The Andrew W. Mellon Foundation Ambrose Monell Foundation The G. Unger Vetlesen Foundation

\$10,000 to \$99,999

Anonymous The Abell Foundation, Inc. Air Liquide Foundation Association of American Medical Colleges The Bodman Foundation The Brinson Foundation Dana and Albert R. Broccoli Charitable Foundation Carnegie Institution of Canada/Institution Carnegie du Canada The Crystal Family Foundation Dow AgroSciences LLC Durland & Co., Inc. The O.P. and W.E. Edwards Foundation Fondation de France Herman Frasch Foundation for Chemical Research Golden Family Foundation Robert and Margaret Hazen Foundation Richard W. Higgins Charitable Foundation Laurel Foundation The McMurtry Family Foundation MGW & CJW 2007 Trust The Robert & Bethany Millard Charitable Foundation Monsanto Company The Kenneth T. and Eileen L. Norris Foundation Northrop Grumman Corporation

The Pfizer Foundation Matching Gifts Program The Rose Hills Foundation Shippy Foundation Skoll Global Threats Fund Society for Developmental Biology Syngenta Biotechnology, Inc. The Sidney J. Weinberg, Jr. Foundation

\$500 to \$9,999

Anonymous American Academy of Arts and Sciences Arnhold Foundation, Inc. Bank of America Matching Gifts The Boeing Gift Matching Program The Bristol-Myers Squibb Foundation, Inc. Carl Zeiss Microscopy, LLC Cavalieri-Look Fund Coulombe Family Trust Damel Investors LLC The Nick Dewolf Foundation Earth Force, Inc. Ernst Charities Arthur and Linda Gelb Charitable Foundation Google Matching Gifts Program Hicks Family Charitable Foundation The Johns Hopkins University The Johns Hopkins University - Department of Biology The Marion I. & Henry J. Knott Foundation, Inc. Laubach Family Fund Linden Trust for Conservation Longfield Family Charitable Foundation Lutron Foundation Robert W. and Gladys S. Meserve Charitable Trust Mulago Foundation Omenn-Darling Family Advised Fund Honey Perkins Family Foundation, Inc. Rathmann Family Foundation **Roxiticus Foundation** T. Rowe Price Foundation, Inc. Harvey and Leslie Wagner Foundation Whittier Trust Company Xerox Corporation ZGF Architects LLP The Zoback Trust

Lifetime Giving Societies

The Carnegie Founders Society



Andrew Carnegie, the founder of the Carnegie Institution, established it with a gift of \$10 million. Although he ultimately gave a total of \$22 million to the institution, his initial \$10 million gift represents a special level of giving. In acknowledgment of the significance of this initial contribution, individuals who support Carnegie's scientific mission with lifetime contributions of \$10 million or more are recognized as members of the Carnegie Founders Society.

Caryl P. Haskins* William R. Hewlett* George P. Mitchell

The Edwin Hubble Society



of work before major discoveries like these can be made. The Edwin Hubble Society honors those whose lifetime support has enabled the institution to continue fostering such long-term, paradigm-changing research by recognizing those who have contributed between \$1,000,000 and \$9,999,999.

D. Euan and Angelica Baird William and Cynthia Gayden Michael and Mary Gellert Robert G. and Alexandra C. Goelet William T. Golden* Crawford H. Greenewalt* David Greenewalt* Margaretta Greenewalt* Robert and Margaret Hazen William R. Hearst III Richard E. Heckert* Kazuo and Asako Inamori Michael T. Long Burton and Deedee McMurtry Jaylee* and Gilbert* Mead Cary Queen Deborah Rose, Ph.D. William J. Rutter Thomas and Mary Urban Sidney J. Weinberg Jr.*

The most famous astronomer of the 20th century, Edwin Hubble, joined the Carnegie Institution

in 1919. Hubble's observations shattered our old concept of the

universe. He proved that the universe is made of collections of galaxies and is not just limited to our own Milky Way—and that it is expanding. This work redefined the science of cosmology. Science typically requires years

Second Century Legacy Society

The Vannevar Bush Society



Vannevar Bush, the renowned leader of American scientific research of his time, served as Carnegie's president from 1939 to 1955. Bush believed in the power of private organizations and wrote in 1950, "It was Andrew Carnegie's conviction that an institution which sought out the unusual scientist, and rendered it possible for him to create to the utmost, would be worthwhile . . ." He further said that "the scientists of the institution

... seek to extend the horizons of man's knowledge of his environment and of himself, in the conviction that it is good for man to know." The Vannevar Bush Society recognizes individuals who have made lifetime contributions of between \$100,000 and \$999,999.

Anonymous (3) Philip H. Abelson* Bruce and Betty Alberts Daniel Belin and Kate Ganz Bradley F. Bennett* Didier and Brigitte Berthelemot Gary P. and Suzann A. Brinson Donald and Linda Brown Richard Buynitzky* A. James Clark Tom and Anne Cori John Diebold* Jean and Leslie Douglas* James Ebert* Bruce W. Ferguson and Heather R. Sandiford Stephen and Janelle Fodor Henrietta W. Hollaender* Antonia Ax:son Johnson and Goran Ennerfelt Paul and Carolyn Kokulis Gerald D. and Doris* Laubach Lawrence H. Linden John D. Macomber Steven L. McKnight Richard A. and Martha R. Meserve J. Irwin Miller* Al and Honey Nashman Evelyn Stefansson Nef* Alexander Pogo* Elizabeth M. Ramsey* Vera and Robert* Rubin Allan R. Sandage* Leonard Searle* Frank Stanton* Christopher and Margaret Stone William and Nancy Turner

The Carnegie Institution is now in its second century of supporting scientific research and discovery. The Second Century Legacy Society recognizes individuals who have remembered, or intend to remember, the Carnegie Institution in their estate plans and those who have supported the institution through other forms of planned giving.

Anonymous (2) Paul A. Armond, Jr. Lore E. Brown Eleanora K. Dalton Nina V. Fedoroff Kirsten H. Gildersleeve Gary K. Hart and Cary S. Hart Robert and Margaret Hazen Paul A. Johnson* Paul and Carolyn Kokulis Gilbert and Karen Levin Chester B. and Barbara C. Martin Robert Metcalf Al and Honey Nashman Holly M. Ruess Leonard Searle* Maxine and Daniel Singer Thomas H. B. Symons, C.C. John R. Thomas, Ph.D. Hatim A. Tyabji

* **Deceased:** Members were qualified with records we believe to be accurate. If there are any questions, please call Irene Stirling at 202.939.1122.

Honors & Transitions

Honors

Trustees/Administration

Carnegie trustee **Sandra Faber** was awarded the 2012 Catherine Wolfe Bruce Gold Medal for lifetime achievement in astronomical research.

Carnegie president **Richard Meserve** was elected a foreign member of the Russian Academy of Sciences. He was also elected president of the Harvard Board of Overseers. Meserve was awarded the William S. Lee Award for Leadership by the Nuclear Energy Institute and the binaugural 2011 Richard L. Garwin Award from the Federation of American Scientists.

Retired director of administration and finance **Gary Kowalczyk** received Carnegie's 2011 Service to Science award.

Embryology

Steve Farber's BioEYES program received the 2012 Viktor Hamburger Outstanding Educator Prize from the Society for Developmental Biology.

Staff associate **Christoph Lepper** received a National Institute of Health Director's Early Independence Award.

Geophysical Laboratory

Robert Hazen received the 2012 Virginia Outstanding Faculty Award from the State Council of Higher Education for Virginia.

New staff member **Timothy Strobel** was awarded the 2011 Jamieson Award by the International Association for the Advancement of High Pressure Science and Technology.

Global Ecology

Greg Asner was named a Senior Energy and Climate Partnership of the Americas Fellow by the U.S. Department of State.

Plant Biology

Director **Wolf Frommer** was awarded the Lawrence Bogorad Award for Excellence in Plant Biology Research by the American Society of Plant Biologists.

Terrestrial Magnetism

Richard Carlson was elected to the National Academy of Sciences.

Librarian **Shaun Hardy** received Carnegie's 2011 Service to Science award.

Transitions

Trustees/Administration

The Carnegie board of trustees welcomed **Michael Long** and **Cristián Samper** as new board members.

Former senior board member **Michael Duffy** rejoined the board as a full member.

Director of administration and finance **Gary Kowalczyk** retired in 2011.

Cynthia Allen became the new director of administration and finance.

Embryology

Christoph Lepper was appointed a staff associate July 1, 2011.

Terrestrial Magnetism

Former department director **Sean Solomon** is taking a leave of absence to serve as the new director of Columbia University's Lamont-Doherty Earth Observatory.

Linda Elkins-Tanton became the new department director on September 26, 2011.

Carnegie Institution for Science

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Friends, Honors & Transitions



★ Sandra Faber



★ Richard Meserve



★ Gary Kowalczyk



★ Steve Farber



★ Christoph Lepper



★ Robert Hazen



 \star Timothy Strobel



★ Greg Asner



★ Wolf Frommer



★ Richard Carlson



★ Shaun Hardy



★ Michael Long



★ Cristián Samper



★ Michael Duffy



 \star Cynthia Allen



★ Sean Solomon



★ Linda Elkins-Tanton

Research Highlights

Embryology

Deciphering the Complexity of Cellular, Developmental, and Genetic Biology



dietary fat increases dietary cholesterol absorption. Researchers know that after cholesterol is absorbed by enterocytes it combines with proteins to form lipoproteins, which distribute lipids throughout the body. A protein called NPC1L1 plays an important, albeit poorly under-

Fish, Fat, and Cholesterol

Balanced cholesterol levels are essential for heart health, but many aspects of cholesterol's transport remain poorly understood. Carnegie's Steven Farber, James Walters, and Jennifer Anderson have developed a strategy that allows scientists to watch lipid metabolism in live zebrafish. They made the first visual observations of cholesterol absorption in a living vertebrate system, enabling studies to better understand diabetes, obesity, and cardiovascular disease.

Fatty acids, cholesterol, and most other lipids are absorbed by the small intestine in vertebrates. The complexity of cells, fluids, microorganisms, and bile make it very difficult to study lipid metabolism outside of the context of the living body.

When we consume fat (triacyglycerides), our digestive organs release enzymes that break it down into fatty acids. Despite years of study, the cellular processes that mediate dietary fatty acid uptake into absorptive cells (called enterocytes) are unclear. Once absorbed, these fatty acids are prepared for transport out of the cell, transformed into droplets of stored fat, or burned by the cell for energy. Many of the steps that regulate the formation of these intracellular fat droplets are unknown. In addition, scientists have yet to explain a long-standing observation that





Clear, living zebrafish gorge on fluorescently tagged chicken egg yoke (top right). In panel two, inside the intestines, the bright green areas are the lipids. Two cholesterol transporter cells are shown in red on right. The bottom panel shows glowing red nuclei that provide a landmark for the scientists, while cholesterol glows green. 28

Embryology

Embryology, Continued



Farber lab's James Walters is part of the team that studies lipid metabolism using zebrafish.

stood, role in cholesterol absorption by the enterocytes.

To address these unanswered questions, the Farber lab turned to the optically clear zebrafish larvae. Using fluorescent forms of lipids to observe fat and cholesterol absorption in the small intestines of live zebrafish, they made a number of breakthroughs. They confirmed that the physiological processes regulating fatty acid and cholesterol absorption are linked, which was first suggested in the 1960s. They also found that a long-chain fatty acid (oleic acid) could increase dietary cholesterol uptake by modulating the subcellular location of NPC1L1. Their research further revealed that-following a meal rich in fat and cholesterol-fatty acids were rapidly stored as lipid droplets while cholesterol was initially stored in special structures called endosomes. These results and this novel research strategy will help us to better understand the cell biology of lipids, with important implications for human health.

Map Reading: Using Proteins to Understand Genes

The ability to sequence genomes has outstripped the research methods for deciphering the information they encode. Carnegie's Nick Ingolia has been working to develop and refine new techniques for mapping protein synthesis, which can be used to decode these genomes. His work is revealing a previously little-known level at which gene expression is controlled and, in the process, helping to home in on potential cancer treatment targets.

The genome of a cell is made up of many thousands of genes. Each gene is an instruction for making one protein. But different genes are activated at different times and in different types of cells. When a gene is turned on, it is *transcribed* into bits of specialized RNA, called mRNA. This mRNA leaves the nucleus and is *translated* into a sequence of amino acids that, when complete, forms a protein. Ingolia's methods reveal many examples where mRNAs are transcribed but where their translation into protein is specially regulated, or occurs in ways that were not previously appreciated.

Only some portions of the genome are used to generate mRNAs. Other portions regulate mRNA translation or serve purposes that aren't yet understood. Ingolia and his colleagues focused on mapping the regions of the mammalian genome that are actually translated into protein. This is different from earlier techniques that used computational modeling to predict the parts of the genome that are turned into protein; Ingolia's team showed those models are incomplete. For example, sometimes the region of a gene that is translated into protein is longer or shorter than expected and the difference between these expectations and actual protein synthesis could change the way the protein functions. Also, sometimes a cell will translate a "decoy"







segment of the genome, which resembles the code for a particular protein, but is different. It serves to distract the translation of the real gene into a protein, thus controlling the amount of protein produced.

Ingolia's team also measured the speed of protein production to gain a better understanding of places where translation stalls or slows down. Lastly, they focused on changes in protein synthesis that take place as mouse embryonic stem cells differentiate into specialized cells that serve specific functions. Stem cells initially divide rapidly into more stem cells. Eventually they differentiate into different types of cells. Ingolia found that translation of suites of genes in the mouse genome changed when the cells went through this process. Interestingly, these same genes were found to be activated in cancer cells and were inhibited by a cancer-fighting drug. □

Top left: Ribosomes attach to mRNA and, based on mRNA instructions, manufacture proteins.

Bottom left: Genes make proteins. First, the information from the bases are copied from a DNA strand into a strand of messenger RNA (mRNA). The mRNA leaves the cell nucleus for an organelle called the ribosome, where it directs the production of amino acids that form the protein.



Embryology's Nick Ingolia

Geophysical Laboratory

Probing Planetary Interiors, Origins, and Extreme States of Matter



Life's Poisonous Start

Carbon, the basis for life, either came from comets and asteroids during the formation of the early Solar System or was trapped within the Earth during planetary accretion. How the Earth was able to hold onto this carbon during its energetic origin is not known. Staff scientist George Cody, postdoctoral fellow Yoko Kebukawa, and their colleagues have been studying organic carbon in comets and meteorites. They find that formaldehyde, a poisonous compound, may be the origin of the complex carbon molecules in these bodies and, by extension, may be the source of the existence of life.

It has long been a challenge to determine how early carbon-containing materials were formed, because their chemistry is extremely complex. Cody and team succeeded in determining their origin by applying advanced molecular spectroscopic techniques both at the Geophysical Laboratory—using solid-state 13 C nuclear magnetic resonance spectroscopy-and at the Advanced Light Source (an X-ray synchrotron radiation source)—using X-ray absorption spectroscopy. Using these techniques, they studied the chemistry and structures of actual particles of carbon-containing material from comet Wild-2, retrieved by the NASA Stardust mission; of interplanetary dust

particles, obtained by high-flying aircraft; and of ancient meteorites called chondrites found on Earth. All contain a polymer (a large molecule) formed by formaldehyde.

Formaldehyde (CH₂O) is a common molecule in the galaxy; radio astronomers observe it in the interstellar medium, in protoplanetary disks, and in cometary



Top: The image at left shows a scanning electron microscope of tiny carbon-containing spheroids made during the synthesis of formaldehyde. The image at right shows material from a typical, ancient meteorite called a chondrite. The matter in red is carbon, blue is iron, and green is calcium.

Bottom: George Cody (left) and Yoko Kebukawa are in their lab.



volatiles. Under certain circumstances, formaldehyde will spontaneously grow into a polymer through a complex series of reactions. Cody and team are now tackling the details of the formaldehyde reaction in the lab. Interestingly, along with the polymerization of formaldehyde, simple sugars are formed—including ribose, a component of ribonucleic acid (RNA), which is essential to life. The polymerization of formaldehyde also yields extremely small spheroidal particles that are similar in size and shape to the nanoglobules in primitive meteorites.

Currently the team is focused on the effect of synthesis temperature on molecular structure and the role of ammonia in the reaction. Ammonia is also abundant in the galaxy and is expected to play an important role in the complex formaldehyde polymerization chemistry. Their results support a new theory that the carbon components in chondrites and comets were formed through the polymerization of interstellar formaldehyde after the accretion of small planetesimals in the presence of water during the early Solar System.

Tiny Cages Hold Big Promise

Understanding the chemical reactions that can create tiny molecular cages that hold other "guest" molecules structures called clathrates—is key to creating a new generation of electronic devices and possible energy materials. Timothy Strobel and team are the first to use high-pressure synthesis to create a stable clathrate of sodium (Na) and silicon (Si)—the least understood system of the so-called group 14 clathrates. Strobel also created a new clathrate of hydrogen sulfide (H_2S) and molecular hydrogen (H_2). Both findings open the door for major advances in materials science.

The sodium silicon structure NaSi₆ shows what is called sp³ bonded silicon atoms (blue) which form tunnels along the a-axis of the crystal, trapping sodium atoms (yellow).



Geophysical Laboratory, Continued

nages courtesy Timothy Strobe



Top: This image is a composite of time-lapse snapshots overlaid for the cage-like H_2S+H_2 clathrate structure at 20,000 times atmospheric pressure (2 GPa). It shows ordered hydrogen bonds for H_2S molecules (yellow), hydrogen atoms (white), and disordered H_2 molecules (red).

Bottom: Timothy Strobel (foreground) works with a supercontinuum light source.

Until now, scientists created most silicon clathrates by heating a chemical precursor in a vacuum. This process, however, is not ideal for controlling growing conditions and keeping the cages stable. Synthesis under high pressure provides a reliable means to control growth of certain other materials (e.g., diamond and cubic boron nitride), so the Strobel team decided to pursue that approach for clathrates. They subjected mixtures of Na-Si to various pressures and temperature regimes and found a type of clathrate, Na₈Si₄₆, that formed at pressures ranging from 20,000 to 60,000 times atmospheric pressure (2 to 6 gigapascals, GPa) and temperatures of 1160 to 1520°F (900 to 1100 K). When the pressure was increased to 79,000 atmospheres (8 GPa), a new clathrate structure NaSi₆ formed. The latter material behaves like a metal and has never been seen before.

In addition to these experiments, the researchers performed calculations to predict how the materials would behave. Calculations and experiments revealed that sodium clathrates are thermodynamically stable at high pressure. The consistency suggests that scientists can use theoretical calculations to predict new synthesis routes for other compounds.

Strobel and colleagues also discovered that a clathrate formed from hydrogen sulfide and molecular hydrogen $(H_2S + H_2)$ exhibited different behaviors under different pressure conditions. At pressures of 35,000 atmospheres (3.5 GPa), the compound crystallized into a stable "guest /host" structure held together by weak attractive forces (van der Waals forces). At 170,000 atmospheres (17 GPa), the hydrogen bonding increased between neighboring H_2S molecules. This type of bonding alone, however, was not sufficient for stability; when the hydrogen molecules were removed from the cage, the H_2S molecule framework collapsed. The research showed that novel cooperative interactions arising from molecular packing are necessary to hold such structures together. \Box

Global Ecology

Linking Ecosystem Processes with Large-scale Impacts



Deciphering Dead Zones

There are some 400 coastal "dead zones" around the world where the concentration of dissolved oxygen is so low that water cannot support healthy aquatic life. Researchers know that nutrients from agricultural runoff play an important role in the formation of dead zones, but their impact relative to other factors—such as the degree of stratification or layering in the water body—is difficult to identify. This makes it challenging to assess how the incidence and size of dead zones can be reduced through agricultural and runoff management.

Dan Obenour, a Ph.D. student working with Anna Michalak and a colleague at the University of Michigan, conducted the first study to isolate the impact of nutrients from that of stratification on dead zone variability in the Gulf of Mexico. They developed a statistical model that can pinpoint the impact of stratification; they found that stratification and nutrient concentrations contribute approximately equally to reductions in dissolved oxygen concentration. This result means that reducing nutrient runoff into the Gulf would have a substantial beneficial effect—a boon to coastal management practices.

Strong stratification, measured by looking at the temperature or salinity as a function of depth, shuts off the resupply of oxygen to the bottom. Nutrients stimulate algal growth, and algae consume oxygen as they sink to the bottom and decompose. The net impact of both factors is a reduction in oxygen levels and the formation of dead zones.

Obenour and team used a data analysis tool called geostatistical regression to isolate the impact of stratification by looking at its spatial relationship to dissolved oxygen at the bottom of the water column. They modeled 10 years of monitoring data from 1998-2007 along the Louisiana-Texas coast. They examined yearly variability in dissolved oxygen that could not be explained by the relationship with stratification against a variety of other factors; they found that nitrogen, in the form of nitrate and nitrite concentrations, explained the remaining signal.

The team then explored the impact of hypothetical nutrient reduction scenarios over the 10 years. They found that shrinking the size of the low-oxygen zone to 1,930



Ph.D. student Dan Obenour is on a research vessel in the Gulf of Mexico study area.

Global Ecology, Continued





square miles (5000 km²), a goal set for the Gulf of Mexico by the Mississippi River/Gulf of Mexico Watershed Nutrient Task Force, would require about a 30% to 60% reduction in nitrate and nitrite concentrations.

Uncovering Canopy Chemistry

In seeking to understand the chemistry of the plant canopy, scientists try to determine the different effects of environmental regulators—such as terrain and soil—and of different plant species. The failure to understand these effects limits our ability to predict the effects of climate change and to unravel how nitrogen and carbon cycle through a system. A study by doctoral student Kyla Dahlin has revealed some new insights about whether the stronger predictors of canopy chemistry are environmental regulators or species within plant communities.

To study the diversity and distribution of plants, researchers can look at how different plant traits, like plant height or leaf shape and size, among different species in a community vary under different environmental conditions. Dahlin took a broader approach. For her Ph.D. thesis, she measured four chemical traits of California plant communities using the new Carnegie Airborne Observatory Airborne Taxonomic Mapping System (AToMS), coupled with on-ground sampling, to produce maps of these four traits across the landscape.

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Global Ecology



Dahlin measured leaf nitrogen per unit mass, leaf carbon per unit mass, leaf water concentration, and canopy water content across the ecologically diverse Jasper Ridge Biological Preserve in Northern California. The plant communities included savanna/grasslands, evergreen oaks and chaparral, riverbank ecosystems, and more. AToMS allowed Dahlin to detect the chemistry at extremely high resolution.

Dahlin combined her airborne data with known environmental regulators, such as terrain, soil, and land-use history, to model whether environmental regulators or species within plant communities were more indicative of the chemical variation of the canopy. She found that environmental regulators played a role but that plant communities were a stronger predictor.

For instance, Dahlin found that leaf shedding (deciduousness) willow-dominated communities, areas dominated by deciduous shrubs, and tracts with deciduous oaks were stronger predictors of leaf nitrogen levels than environmental regulators, making plant characteristics more important than environmental ones. Since nitrogen and carbon variations in the canopy are more closely tied to the vegetation community than to environmental regulators, individual species could become critical for understanding nitrogen and carbon cycling. The results not only are intriguing, they also represent a major step in integrating remote-sensing science with community ecology and spatial statistics. □



Top: The left image shows a true color image of part of Jasper Ridge Biological Preserve. The right image shows the same area but colored by canopy chemistry measured by the Carnegie Airborne Observatory Airborne Taxonomic Mapping System. Leaf carbon is red, leaf nitrogen is green, and leaf water is blue.

Bottom: Kyla Dahlin assesses the health of a young California bay laurel *(Umbellularia californica).*


Observatories

Investigating the Birth, Structure, and Fate of the Universe



Birthing Stars

The birth of a star is a violent business. It begins in giant clouds of molecular gas and dust that gravitationally collapse. Pressure and temperature increase, causing these clumps to coalesce into spinning spheres of very hot gas known as protostars. Over millions of years the protostars grow and ultimately stabilize into different star types, depending on their mass. Miguel Roth and team are studying whether young stars of different masses are made by the same mechanisms. Low-mass stars can eject energetic jets of gas in an early stage, exciting the surrounding medium. The excitation induces chemical changes and forms spectacular knots of molecular hydrogen and sulfur called Herbig-Haro objects in the neighborhood of the newborn star. These knots are short-lived, compact nebula with specific chemical signatures. Roth and team believe that such knots could be associated with the formation of more massive stars as well.

Other processes can excite hydrogen, too. In a recent study, the team looked at star clusters of intermediate mass stars—stars about 7 to 9 times the mass of the Sun—in the Carina Nebula. The group consisted of at least 72 young stars in a cloud of ionized hydrogen called an H II region.

The astronomers found a photodissociation region caused by ultraviolet (UV) photons from other, more massive stars in the area, plus stellar winds that provide the heat that affects the gas chemistry. They found the 72-

The inset of this image shows the area (in red) where Miguel Roth and colleagues discovered the 72-member group of young stars in the Carina Nebula. They are embedded in filaments of a hydrogen cloud, a typical birthplace for newborn stars.







member cluster of young stars embedded at the top of a pillar-like structure made of dust just behind this region.

The scientists discovered that their young star group shares many features of recent star formation with stars of different masses formed elsewhere in the Carina Nebula. Scientists believe that the heating mechanism of star formation for those other stars was the interaction of UV radiation from a previous generation of stars and winds with the interstellar medium. They think that the original molecular cloud interacts with this radiation and forms photodissociation areas, which appear as dust structures that resemble pillars and "elephant trunks." Stars form toward their tops. Roth and team believe that the same formation scenario is likely for their new group of 72 stars. The astronomers hope to unravel the different formation processes with a new infrared spectrograph on the Magellan telescope called FIRE, an instrument built by investigators at MIT.



Miguel Roth (back) and his colleagues Paolo Persi (middle) and Mauricio Tapia are on a three-night observation run.

Behold Galacticus!

Galacticus is not a super hero. Galacticus is a super model used to determine the formation and evolution of the galaxies. Developed by Andrew Benson, the George Ellery Hale Distinguished Scholar in Theoretical Astrophysics, it is one of the most advanced models of galaxy formation available.

Rather than building his model around observational data, Benson's Galacticus relies on known laws of physics and the so-called N-body problem, which predicts the motions of celestial bodies that interact gravitationally in groups. Galacticus' results are stunning 3-D videos.

Some 80% of the matter in the universe cannot be seen. This unseen matter is believed to be cold, dark matter, and it forms a halo around galaxies like our Milky Way. Dark matter affects how galaxies form. Galacticus solves the physics of galaxy formation inside a hierarchy of dark matter halos, which are linked in tree-like structures called merger trees. Galacticus is "fed" merger trees and then populates them with galaxies.

Benson and colleagues have run numerous simulations, including a simulation of the reionization period—an epoch between about 200 million and one billion years after the Big Bang—when neutral hydrogen began to form quasars, stars, and the first galaxies. Benson examined the growth of ionization fronts—clouds of hydrogen in the intergalactic medium that has been stripped of electrons. He modeled a star-forming protogalaxy with an energetic, dense, bright area near its center called an Active Galactic Nucleus (AGN). The stars and AGN yielded ultraviolet and X-ray photons that produced an ionization front. The results will guide future surveys of this epoch.

Benson, with colleague Arif Babul, is also modeling the growth of black holes with highly energetic jets at their

Observatories, *Continued*

courtesy Andrew Be



center within a hierarchically growing population of galaxies. Researchers believe the spin of the black hole drives the jets; the model computes this spin rate and the power of the jets.

Benson has ambitious plans for Galacticus, including simulating a very early and hidden part of the universe dubbed the submillimeter universe when light was obscured by dust. Thus far, theorists have not been able to model it successfully. \Box

Left top: This still image shows a representation of a merger tree from one of Galacticus' calculations. Each colored sphere represents a dark matter halo that grows with time and eventually merges with other dark matter halos.

Below: Andrew Benson at the Griffith Observatory



Right: This image shows a simulated patch of sky containing many thousands of galaxies. Benson's model produced this image; it represents what the Hubble Space Telescope would see if it were able to "observe" Benson's model universe. It was made with collaborators Rick White and Mike Fall.



Plant Biology

Characterizing the Genes of Plant Growth and Development



Pumped Up Plants

Steroids are important hormones in both animals and plants. Steroids bulk up plants just as they do human athletes, but the pathway of molecular signals that tell the genes to boost growth and development is more complex in plant cells than in human and animal cells. Unlike animals, plants do not have glands to produce and secrete hormones. Rather, each plant cell has the ability to generate hormones. Another difference is that animal cells typically have receptor molecules that respond to select steroids located within a cell's nucleus. In plants, steroid receptors are anchored to the outside surface of a cell's outer membrane—the membrane that delineates a cell as a single unit thus requiring relay systems to trigger changes in gene regulation in the nucleus.

In a tour de force, Carnegie's Zhiyong Wang has been homing in on the chemical signaling pathways of one major class of plant hormones called brassinosteroids, making it one of the best-studied aspects of cellular physiology. As it turns out, brassinosteroids are not only involved in pumping up the plant, they conduct in an incredibly wide array of functions, including response to environmental stresses, cell elongation, and resistance to pathogens.

Mutant plants that are deficient in brassinosteroids show many defects at various phases of the plant life cycle, including reduced seed germination, irregular growth in the dark, altered pattern of organ formation, dwarfism, and sterility. Brassinosteroids also control the leaf angle, which is critical for optimal energy capture and thus yield. As such, understanding the brassinosteroid pathway could help researchers improve plant growth and hardiness, which could increase crop yields and help fight world hunger.

The Wang lab has become the leader in uncovering this exciting signaling pathway. In 2012, Wang and his lab again expanded our understanding of many important aspects of the pathway. They made new breakthroughs in



A plant's leaves are sealed with a gastight wax layer to prevent water loss. A plant breathes through microscopic pores called stomata (Greek, for mouths) on the surfaces of its leaves, shown here. Zhiyong Wang and team found that brassinosteroids are involved in regulating the developmental process for these crucial plant organs.

Plant Biology, Continued





When a seed germinates underground, it rapidly elongates its stem to reach the soil's surface. If its neighbor shades the plant, it elongates its stem to out compete its neighbors for sunlight. Under full sunlight, the plant's priority is leaf expansion. Management of these responses to environmental signals is, at least in part, controlled by a "command center" that involves brassinosteroids and another major class of plant hormones called gibberellins.

how brassinosteroids are involved in the distribution of a plant's gas-exchange system, as well as breakthroughs in how a plant responds to changes in light and temperature. Moreover, they discovered a system of "cross-talk," by which brassinosteroids interact with a chemical signaling system controlled by another major class of plant hormones called gibberellins. Together, brassinosteroids and gibberellins form a "command center" controlling plant growth and environmental responses. Wang believes this network will be a major target for genetically engineering high-yielding crops.

A Paradigm Shift in Plant Disease Resistance

Feeding a growing population, resisting plant disease, and overcoming the human problems of obesity and diabetes have one thing in common: they involve sugar transport by SWEET proteins. Plants and humans use sugars as key compounds to distribute carbon, energy, and a common set of transport proteins (SWEETs) that Wolf Frommer's lab discovered. SWEETs hold the key to increasing crop yields. They open a completely new way to prevent plant diseases. Excitingly and unexpectedly, they also may help researchers understand how sugars are transported in human cells, with implications for our understanding of diabetes and obesity.

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Plant Biology



Plants capture light energy and convert atmospheric carbon dioxide and water to create the sugar sucrose. Sucrose biosynthesis occurs inside green leaf cells, which are connected to all other plant organs via a network of veins, the phloem. Unlike animals, plants do not have a mechanical pump for distributing sugars and other nutrients; they use a molecular pump instead. Over the course of 20 years, Wolf Frommer and his lab have identified the core components of this pump—starting with identifying the active pore proteins that load sucrose into the plant's veins. But the mechanism that gets the sugar out of the cells that produced it in the first place remained elusive.

Frommer's team tracked down these missing links, named SWEETs, with novel tools: fluorescent Förster resonance energy transfer biosensors. These sensors report sugar levels in subcompartments of individual cells and quantify them by a color change. The technology has myriad research uses, including the study of brain chemistry, cancer, and biofuel development.

Over the past two years, the team discovered that bacteria and fungi hijack the SWEET sugar transport pores to access the plant sugar. The pathogens enter the spaces between the cells, camp out, pick the lock to the pantry, dine, reproduce, and dash. The pathogens inject activator proteins into the cell that directly turn on sugar efflux transporters—the floodgate mechanism at the plasma membrane that pumps sugars out of the cell for the pathogen to steal. This is the first time scientists have a direct handle on controlling the food supply to pathogens, which opens a new avenue for preventing a wide range of crop diseases. Additionally, fine-tuning the SWEETs may help produce higher crop yields. And because humans

> also use SWEETs to transport glucose, understanding the sugar-transporting mechanism could lead to breakthroughs in medicine. □

Plant Biology director Wolf Frommer (left) discusses the new microfluidic chip for live imaging of plant roots with Li-Qing Chen (center), lead author of the *Science* and *Nature* articles describing the identification of the SWEET transporters, and Guido Grossmann (right), lead author of the *Plant Cell* and *JoVE* articles describing the development and use of the novel RootChip technology.



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Terrestrial Magnetism

Terrestrial Magnetism

Understanding Earth, Other Planets, and Their Place in the Cosmos



Moon Awash with Water

The Moon has much more water than previously thought, according to a team of scientists led by geochemist Erik Hauri. Their research shows that droplets of magma trapped within crystals collected during the Apollo 17 mission contain 100 times more water than earlier measurements recorded. The lunar magmas are just as wet as Earth's mid-ocean-ridge basalts—a common volcanic rock—which means that the Earth's mantle and the Moon's mantle could have similar amounts of water. The finding could overturn the prevailing theory about the Moon's origin.

The team used Carnegie's state-of-the-art NanoSIMS 50L ion probe. This instrument shoots a beam of charged particles at tiny samples to determine their trace element composition with unparalleled precision; the team's unique techniques can detect extremely minute quantities of water in glasses and minerals. The researchers measured seven bits of magma trapped within lunar crystals. Because these bits, called inclusions, are encased in crystals, water and other easily evaporative (i.e., volatile) materials could not escape during the eruptions that brought them to the Moon's surface eons ago.

The Earth and the other inner planets contain relatively low amounts of water and volatile elements, which were





not abundant in the inner Solar System during planet formation. The Moon has even lower quantites of these volatile elements, which has long been claimed as evidence that the Moon must have formed from material ejected from the Earth as a result of a catastrophic giant impact. But this new research shows that aspects of this theory must be reevaluated, because such an impact would have been expected to cause the escape of these volatiles.

The study also puts a new twist on the origin of water ice detected in craters at the lunar poles by several recent NASA missions. The ice has been attributed to comet and meteoroid impacts, but it is possible that some of this ice could have come from the water released by past eruptions of lunar magmas.

Researchers should take these findings into account when analyzing samples from other planetary bodies in our Solar System. Hauri's team believes that their unique method of analysis provides an accurate and direct way to determine the water content of a planet's interior.

Left bottom: Carnegie's Erik Hauri (left) sits with coauthor Alberto Saal in the lab.

Particles to Planets

Comets and asteroids represent the building blocks of our Solar System and thus record the physics and chemistry of how planets formed. But there are puzzles: How did icy comets obtain particles made by high-temperature conditions? How did the particles acquire different compositional layers ("rims"), and when did they form? Collaboration between theoretical astrophysicist Alan Boss and cosmochemist Conel Alexander, with colleague Morris Podolak, is answering these riddles and illuminating the earliest stages of Solar System formation.

The protosun in the early solar nebula is thought to have experienced a series of outbursts propelling particles of many sizes through the nebula to distances ranging from one astronomical unit (the Earth/Sun distance, AU) to 10 AUs. Boss and Alexander are the first to model the particles' trajectories. They modeled 200 2-centimeter spherical particles (representing calcium-aluminum-rich inclusions (CAIs) from primitive meteorites) over 200 years and predicted their thermal and chemical processing.

The model assumed a solar nebula with a mass about 5% of today's Sun, with temperatures ranging from a frigid -351°F (60 K) in the outer regions to a scorching 2240°F (1500 K) near the center. The calculations allowed the CAIs—dominated by the mineral melilite—to orbit, interact with the disk gas and pressure, and gravitationally interact with the protosun.

The 2-centimeter particles started orbiting in unison, but after about 20 model years their trajectories started to diverge significantly. Some particles reached the hot, central disk, while others headed for the cold, outer disk; some did both.

The researchers modeled a number of different evaporation and condensation scenarios for melilite. The model

Left top: This is an optical photograph of a droplet of solidified melt within an olivine crystal from the Apollo 17 mission. It is about 40 microns in diameter. Analysis of the Apollo 17 samples indicates that the Moon has 100 times more water than found in previous measurements.

Terrestrial Magnetism, Continued



indicated that the particles could have undergone processing in the hot inner disk and then moved to the frigid outer regions, explaining how particles with compositions indicating high temperatures of formation could have ended up in icy comets. The results also showed some strange trips back and forth, which could explain the layers with different oxygen isotopes found in some particles. (Isotopes are varieties of oxygen with different numbers of neutrons.) The varying isotope compositions reflect different processing conditions: some isotopes are associated with processing near the protosun, while others are associated with the planet-formation areas. The fact that particles survived the mayhem suggests that they were processed during a late outburst, which counters current assumptions that they formed during earlier stages of Solar System formation. \Box



Top left: This X-ray image shows the compositional differences along the rim of a typical Type A calcium-aluminumrich particle. This particle was found within the Allende meteorite.

Bottom left: Cosmochemist Conel Alexander (left) and theoretical astrophysicist Alan Boss (right) collaborated on modeling the trajectories of tiny particles in the early Solar System to decipher what happened during the earliest stages of Solar System formation.

First Light & The Carnegie Academy for Science Education

Teaching the Art of Teaching Science and Math



Retooling Teaching With Friendship

This past year, Carnegie Academy for Science Education (CASE) staff (Julie Edmonds, Toby Horn, and Marlena Jones) teamed with Friendship Collegiate Academy, a public charter school in northeast Washington, D.C., to help develop a new teaching approach in the school's Allied Health and PolyTech Career Academies. Friendship's objective is to emphasize interdisciplinary Science, Technology, Engineering, and Math (STEM) instruction integrated with English and social studies across its academies.

Seventeen 11th and 12th grade teachers participated in the CASE Summer Institute to learn how to teach using project-based learning. In this approach, the teacher becomes an academic coach and guides students to take control of a topic within a theme and explore it independently.

CASE encouraged the interdisciplinary teams from each academy to choose an interesting, challenging, and rich question to drive instruction across the disciplines. In the Allied Health project, "What is Healthy Living?," a social studies segment examines historical farming practices in the United States, including foods and recipes. The science unit investigates the physics behind Olympic gymnastics, while the microbiology unit explores the role of microbes in food and disease. Other courses contribute similarly relevant topics. The academy's effort will culminate in a community health fair and walk, with student projects on display.

The PolyTech Academy chose to examine the Washington, D.C., subway system, Metro. Facets of this project, "Is Metro Safe, Reliable, and Effective?," are exploring the history of transportation in the city, how to redesign cars to hold more passengers without them feeling crowded, and the physics behind the Metro accident that killed nine people near the Fort Totten station in 2009. At the end of this project, the students will present their recommendations for improvements to the Metro Board.

Some of the teachers were also blended into a six-week student research project in astrobiology, "Life in Extreme Environments," led by Julie Edmonds. Since most of the students in the program attend Friendship, the teachers could see firsthand how their own students became engaged when given the chance to learn independently.

Math for America Spreads its Wings

This past academic year 14 Math for America (M*f*A) fellows taught mathematics to some 1,200 high-need students in Washington, D.C., public schools. Six more fellows began teaching in the current school year, and four more began this year's 15-month training program. Carnegie and American University run the Washington, D.C., M*f*A (M*f*A DC) program.

All MfA fellows attend professional development sessions and are guided by mentors to sharpen their skills.

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First Light & The Carnegie Academy for Science Education

First Light & The Carnegie Academy for Science Education, Continued

Top: PolyTech Academy teachers explore the engineering, physics, and math connections of a proposed student activity before making their decision to pursue the Metro project. From left to right: Jaclyn Claiborne, engineering, math, and statistics; Christian Schaefer, engineering; Aaron Morton, digital graphics and software applications.

Bottom: John Neral, guest speaker from the District of Columbia's Office of the State Superintendent for Education, spoke about the new Common Core State Standards for Mathematics. Seated from left to right: Julie Edmonds, CASE co-director (in yellow); Sarah Bax, CASE Mentor Teacher and MfA DC Master Teacher (in pink); Carlyle Chalmers, Friendship PolyTech Academy video production teacher (in purple); and Francis Ayissi, 10th grade chemistry teacher (in blue).





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In addition to a background in mathematics, teachers must be able to connect with high-need youngsters, have a sense of humor, be flexible, and be resilient. Perhaps most importantly, they need to convey the joy of learning math with a contagious enthusiasm for the subject.

The staff at MfA has made a number of enhancements since 2009, including improved professional development and the inauguration of a Master Teacher Fellowship. James Tanton, who now leads professional development for MfA DC, mesmerizes students and teachers with his animated teaching approach. Tanton deconstructs problems so others understand the reasons why, for instance, multiplying two negative numbers creates a positive one. (See Tanton's style and approach to mathematics at www.jamestanton.com.)

The year 2012 marks the second year for the Master Teacher Fellowship. The goal is to establish five master teachers to encourage math teachers to continue in their field and to help develop leadership skills for high-needs instruction. Master teachers have degrees in mathematics and have taught for at least four years. These fellows receive stipends and professional development support. The first MfA DC Master Teacher, Sarah Bax, received the 2011 Presidential Award for Excellence in Mathematics and Science Teaching. \Box



Top: Mathematician James Tanton directs professional development for the Washington, D.C., Math for America chapter. Tanton has a Ph.D. in mathematics from Princeton University.

Bottom left: Sarah Bax, the first MfA Master Teacher from the Washington, D.C., chapter, received the 2011 Presidential Award for **Excellence** in Mathematics and Science Teaching. She has been teaching in Washington, D.C., public schools for 18 years; she currently teaches at Hardy Middle School.



Financial Profile

for the year ending June 30, 2012 (unaudited)

Carnegie Institution for Science

Reader's Note: In this section, we present summary financial information that is unaudited. Each year the Carnegie Institution, through the Audit committee of its Board of Trustees, engages an independent auditor to express an opinion about the financial statements and the financial position of the institution. The complete audited financial statements are made available on the institution's website at www.CarnegieScience.edu.

The Carnegie Institution of Washington completed fiscal year 2012 in sound financial condition due to the positive returns (+5.30%) of the diversified investments within its endowment; a disciplined spending policy that balances today's needs with the long-term requirements of the institution and the interests of future scientists; and the continued support of organizations and individuals who recognize the value of nurturing basic science.

The single primary source of support for the institution's activities continues to be its endowment. This reliance on institutional funding provides an important degree of independence in the research activities of the institution's scientists.

As of June 30, 2012, the endowment was valued at \$795 million. Over the period 2001-2012, average annual increases in endowment contributions to the budget were 5.5%. Carnegie closely controls expenses in order to ensure the continuation of a healthy scientific enterprise.

For a number of years, under the direction of the Finance committee of the board, Carnegie's endowment has been allocated among a broad spectrum of asset classes including: fixed-income instruments (bonds), equities (stocks), absolute return investments, real estate partnerships, private equity, and natural resources partnerships. The goal of this diversified approach is to generate attractive overall performance and minimize the volatility that would exist in a less diversified portfolio.

The Finance committee of the board regularly examines the asset allocation of the endowment and readjusts the allocation, as appropriate. The institution relies upon external managers and partnerships to conduct the investment activities, and it employs a commercial bank to maintain custody. The following chart shows the allocation of the institution's endowment among asset classes as of June 30, 2012.

Asset Class	Target	Actual
Common Stock	37.5%	40.5%
Alternative Assets	55.0%	54.0%
Fixed Income and Cash	7.5%	5.5%

Carnegie's investment goals are to provide high levels of current support to the institution and to maintain the long-term spending power of its endowment. The success of Carnegie's investment strategy is illustrated in the following figure that compares, for a hypothetical investment of \$100 million, Carnegie's investment returns with the average returns for all educational institutions for the last twelve years.

Carnegie has pursued a long-term policy of controlling its spending rate, bringing the budgeted rate down in a gradual fashion from 6+ % in 1992 to 5.00% today. Carnegie employs what is known as a 70/30 hybrid spending rule. That is, the amount available from the endowment in any year is made up of 70% of the previous year's budget, adjusted for inflation, and 30% of the most recently completed year-end endowment value, multiplied by the spending rate of 5.0% and adjusted for inflation and for debt. This method reduces volatility from year-to-year. The following figure depicts actual spending as a percentage of ending market value for the last 20 years.

In fiscal year 2012, Carnegie benefitted from continuing federal support. Carnegie's federal support has grown from \$24.5 million in 2006 to more than \$34.5 million in new grants in 2012. This is a testament to the high quality of Carnegie scientists and their ability to compete successfully for federal funds in this period of fiscal restraint.

Carnegie also benefits from generous support from foundations and individuals. Funding from foundations has grown from an average of about \$3 million/year in the period from 2000 to 2004 to \$7.4 million in 2012. Within Carnegie's endowment, there are a number of "funds" that provide support either in a general way or targeted to a specific purpose. The largest of these is the Andrew Carnegie Fund, begun with the original gift of \$10 million. Mr. Carnegie later made additional gifts totaling another \$12 million during his lifetime. This tradition of generous support for Carnegie's scientific mission has continued throughout our history and a list of donors in fiscal year 2012 appears in an earlier section of this year book. In addition, Carnegie receives important federal and private grants for specific research purposes, including support from the Howard Hughes Medical Institute for researchers at the Department of Embryology.

Carnegie Institution for Science



Illustration of \$100 Million Investment - Carnegie Returns vs. Average Returns for All Educational Institutions (2000-2012)

Average returns for educational institutions are taken from Commonfund reports on endowment performance.



Endowment Spending as a Percent of Ending Endowment Value*

Statements of Financial Position (unaudited)

June 30, 2012, and 2011

	2012	2011
Assets Current assets: Cash and cash equivalents Accrued investment income Contributions receivable Accounts receivable and other assets Bond proceeds held by Trustee	\$ 2,224,055 47,721 18,495,658 21,436,261 15,694	\$ 1,518,067 0 7,298,027 17,279,764 17,694
Total current assets	\$ 42,219,389	\$ 26,113,552
Noncurrent assets: Investments Property and equipment, net	794,835,568 152,340,983	795,672,507 154,768,137
Total noncurrent assets	\$947,176,551	\$950,440,644
Total assets	\$989,395,940	\$976,554,196
Liabilities and Net Assets Accounts payable and accrued expenses Amount held for others Deferred revenues Bonds payable Accrued postretirement benefits	\$ 11,449,485 0 29,670,190 65,706,919 19,991,999	\$ 10,918,845 0 31,307,772 65,728,416 17,206,079
Total liabilities	\$126,818,593	\$125,161,112
Net assets Unrestricted Temporarily restricted Permanently restricted	\$253,993,414 553,628,669 54,955,264	\$244,949,855 551,513,903 54,929,326
Total net assets	\$862,577,347	851,393,084
Total liabilities and net assets	\$989,395,940	\$976,554,196

Statements of Activities' (unaudited)

Periods ended June 30, 2012, and 2011

	2012	2011
Revenue and support: Grants and contracts Contributions, gifts Other income	\$ 40,529,751 26,801,795 7,820,546	\$ 40,480,694 13,081,598 18,583
Net external revenue	\$ 75,152,092	\$ 53,580,875
Investment income and unrealized gains (losses)	\$ 36,181,149	\$153,142,080
Total revenues, gains, other support	\$111,333,241	\$206,722,955
Program and supporting services: Terrestrial Magnetism Observatories Geophysical Laboratory Embryology Plant Biology Global Ecology Other programs Administration and general expenses	<pre>\$ 14,972,184 20,071,881 20,425,062 11,467,512 10,778,313 8,241,999 852,665 11,259,427</pre>	<pre>\$ 11,957,202 21,920,605 19,962,665 9,670,782 10,032,715 6,267,032 1,050,387 10,556,630</pre>
Total expenses	\$ 98,069,043	\$ 91,418,018
Change in net assets before pension related changes Pension related Changes Net assets at the beginning of the period	\$13,264,198 (2,079,935) \$851,393,084	\$115,304,937 (1,359,592) \$737,447,739
Net assets at the end of the period	\$862,577,347	\$851,393,084

¹Includes restricted, temporarily restricted, and permanently restricted revenues, gains, and other support.

2012 Expenses by Department (\$98.1 Million)



Personnel

July 1, 2011-June 30, 2012

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- ⁶ From May 15, 2012
- ⁷ To March 30, 2012
- ⁸ To August 1, 2011
- ⁹ From June 14, 2012
- ¹⁰ To October 31, 2011
- ¹¹ From April 23, 2012
- ¹² From May 7, 2012
- ¹³ From December 16, 2011
- ¹⁴ From June 9, 2011, to November 30, 2011
- ¹⁵ From September 14, 2011

Carnegie Academy for Science Education

 ¹ From September 1, 2011
 ² From May 31, 2012
 ³ Summer 2011
 ⁴ From April 16, 2012

¹ From April 2, 2012

² From April 25, 2012

³ From August 1, 2011, to October 10, 2011

⁴ From September 10, 2011, to November 20, 2011

⁵ From July 11, 2011, to November 30, 2011



EMBRYOLOGY Front row (left to right): Allan Spradling, Alex Bortvin, Marnie Halpern, Joe Gall, Steve Farber, Chen-Ming Fan, Christoph Lepper, Jeff Han, Fred Tan. Second row: Rafael Villagaray, Ming-Chia Lee, Lynne Hugendubler, Gabriela Rodriguez, Alexis Marianes, Patricia Cammon, Ella Jackson, Rejeanne Juste, Allen Strause, Sean Watson, Earl Potts. Third row: Xiaobin Zheng, Mahmud Siddiqi, Michael Sepanski, Chandra Harvey, Lucy Morris, Jun Wei Pek, Steven Ching, Rob Vary, Vanessa Quinlivan-Repasi, SiewHui Low, Lydia Li, Wilber Ramos, Eric Mills, Tom McDonaugh. Fourth row: Haiyang Chen, Troy Horn, Mohammed Shamim, Yuxuan Guo, Michael Harris, Svetlana Deryusheva, Chun Dong, Oni Mapp, Blake Caldwell, Sara Roberson, Rebecca Obniski, Eric Duboué, Aaron Katrikh, Sheryl Murray, Ethan Greenblatt. Fifth row: Eugenia Dikovskaia, Valeriya Gaysinskaya, Erin Zeituni, Alice Hung, Ona Martin, Dianne Williams, Vicki Losick, Elim Hong, Lei Lei, Vanessa Matos-Cruz, Jianjun Sun, Estela Monge, Ivana Celic, Michelle Macurak, Tagide deCarvalho, Allison Pinder, Yihan Wan, Abhignya Subedi, Glenese Johnson. Sixth row: Zhonghua Liu, Anna McGeachy, Shusheng Wang, Mario Izaguirre-Sierra, Matthew Sieber, Juliana Carten, Yue Zheng, Jen Anderson, Carol Davenport, Christine Pratt, Helan Xiao, Liangji Li, Reid Woods, Diana Camerota. Seventh row: Matthew Brown, Dolly Chin, Micah Webster, Axel Horn, Shiying Jin, Marlow Minor, Pavol Genzor, Irena Martirosyan.

Embryology

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Zheng-an Wu, Special Investigator, NIH Grant (Gall)
Helan Xiao, Carnegie Fellow
Erin Zeituni, Carnegie Collaborative Fellow
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¹ From September 1, 2011 ² From January 23, 2012 ³ From September 22, 2011 ⁴ From November 1, 2011 ⁵ To August 31, 2011; Staff Scientist from September 1, 2011 ⁶ From March 1, 2012 ⁷ To July 15, 2011 ⁸ To February 7, 2012 ⁹ To January 23, 2012 ¹⁰ To August 5, 2011 ¹¹ From June 25, 2012 ¹² To May 31, 2012 ¹³ From August 3, 2011 ¹⁴ From May 1, 2012 ¹⁵ From January 17, 2012 ¹⁶ From March 1, 2012 ¹⁷ To September 9, 2011 ¹⁸ From October 3, 2011 ¹⁹ From November 1, 2011

²⁰ From May 16, 2012 ²¹ To January 13, 2012 ²² To July 31, 2011 ²³ From September 21, 2011 24 From September 12, 2011 ²⁵ To March 23, 2012 ²⁶ To January 27, 2012 ²⁷ To March 31, 2012 ²⁸ To August 31, 2011 ²⁹ From August 4, 2011 ³⁰ To June 30, 2012 ³¹ From October 17, 2011 ³² From October 26, 2011 ³³ From September 12, 2011 ³⁴ To December 5, 2011 ³⁵ From July 1, 2011, to February 29, 2012; JSPS Fellow from March 1, 2012 ³⁶ From November 1, 2011 ³⁷ To November 4, 2011 ³⁸ From April 23, 2012

- ³⁹ From September 1, 2011 ⁴⁰ From July 1, 2011 ⁴¹ To April 29, 2012 ⁴² From September 1, 2011 ⁴³ To December 31, 2011; Research Scientist from January 1, 2012 ⁴⁴ To December 31, 2011 ⁴⁵ From December 12, 2011 46 From July 6, 2011, to August 5, 2011 ⁴⁷ To August 12, 2011 ⁴⁸ To January 10, 2012 ⁴⁹ To August 5, 2011 ⁵⁰ From July 12, 2011, to August 12, 2011 ⁵¹ To May 1, 2012 ⁵² From June 5, 2012 ⁵³ From June 4, 2012 to June 8, 2012 ⁵⁴ From August 23, 2011, to September 16, 2011; From December 16, 2011, to December 31, 2011; From June 13, 2012 55 To August 19, 2011
 - 56 From July 27, 2011, to August 17, 2011 57 To July 2011 ⁵⁸ From July 27, 2011, to August 17, 2011 ⁵⁹ To August 12, 2011 ⁶⁰ From June 9, 2012 61 From June 12, 2012 62 To August 1, 2011 63 Joint appointment with DTM 65 Joint appointment with DTM 66 Joint appointment with DTM 67 From August 4, 2011 68 Joint appointment with DTM 69 Joint appointment with DTM ⁷⁰ Joint appointment with DTM ⁷¹ To July 14, 2011 72 Joint appointment with DTM 73 Joint appointment with DTM
 - ⁷⁴ To May 30, 2012

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- ¹ From March 16, 2012
 ² To March 31, 2012
 ³ From September 1, 2011
 ⁴ From October 3, 2011
 ⁵ From September 1, 2011, to November 5, 2011
 ⁶ To February 19, 2012
 ⁷ To August 31, 2011
 ⁸ From July 1, 2011
 ⁹ From February 28, 2011, to September 15, 2011
- ¹⁰ From July 8, 2011
 ¹¹ To June 30, 2011
 ¹² From September 15, 2011
 ¹³ From July 1, 2011
 ¹⁴ From August 31, 2011
 ¹⁵ From Arguit 16, 2012
 ¹⁷ From February 26, 2012
 ¹⁸ To April 27, 2012
- ¹⁹ From April 16, 2012
 ²⁰ From September 16, 2011
 ²¹ To November 30, 2011
 ²² From June 1, 2012
 ²³ From June 25, 2012
 ²⁴ To November 18, 2011
 ²⁵ To February 15, 2012
 ²⁶ To August 31, 2011
 ²⁷ From September 1, 2011
- ²⁸ From November 1, 2011, to February 27, 2012
 ²⁹ To July 31, 2011
- ³⁰ From November 16, 2011
- ³¹ From December 1, 2011
- ³² From January 16, 2012, to June 22, 2012

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¹ From June 15, 2012 ² From October, 1, 2011 ³ From September 1, 2011 ⁴ To December 22, 2011 ⁵ To September 14, 2011 ⁶ To September 30, 2011 ⁷ To October 31, 2011 8 To August 31, 2011 ⁹ To November 30, 2011 ¹⁰ From May 1, 2012 ¹¹ To August 3, 2011 ¹² To July 20, 2011 ¹³ To June 1, 2012 14 To December 31, 2011 ¹⁵ To June 3, 2012 16 From July 11, 2011 ¹⁷ From January 1, 2012 ¹⁸ From October 1, 2010; not reported previously ¹⁹ From December 6, 2011 ²⁰ To May 31, 2011 ²¹ From April 2, 2012 ²² To September 30, 2011; GMT Site Testing Support ²³ From October 15, 2011 ²⁴ To December 18, 2011 ²⁵ From June 7, 2011; not reported previously ²⁶ To August 29, 2011



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Plant Biology

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¹ From July 2, 2011

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William E. Holt, State University of New York, Stony Brook Emilie E. E. Hooft Toomey, University of Oregon Ya-Ju Hsu, Academia Sinica Dmitri Ionov, Université Jean Monnet, St. Etienne* Neng Jiang, Chinese Academy of Sciences* Catherine L. Johnson, University of British Columbia Karl Kehm, Washington College Katherine A. Kelley, University of Rhode Island Christopher R. Kincaid, University of Rhode Island Carolina Lithgow-Bertelloni, University College London Maureen D. Long, Yale University Mercedes López-Morales, Institut de Ciències de L'Espai (CSIC-ICE) Fukashi Maeno, The University of Tokyo Patrick J. McGovern, Lunar and Planetary Institute Wendy Nelson, University of Houston Francis Nimmo, University of California, Santa Cruz* Fenglin Niu, Rice University Jonathan O'Neil, Laboratoire Magmas et Volcans, Université Blaise Pascal* Morris Podolak, Tel Aviv University Stephen H. Richardson, University of Cape Town Joseph Rodriguez, George Mason University Thomas G. Ruedas Paul A. Rydelek* Alberto Saal, Brown University Brian Savage, University of Rhode Island Martha K. Savage, Victoria University, New Zealand Manuel Schilling, Universidad de Chile* Nicholas Schmerr, Goddard Space Flight Center* Maria Schönbächler, University of Manchester A. M. Celâl engör, Istanbul Technical University* Alison M. Shaw, Woods Hole Oceanographic Institution Yang Shen, University of Rhode Island David W. Simpson, Incorporated Research Institutions for Seismology J. Arthur Snoke, Virginia Polytechnic Institute and State University

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 ¹ From September 26, 2011

 ² To September 25, 2011, then Staff Member

 ³ To December 18, 2011

 ⁴ From September 1, 2011

 ⁵ To September 1, 2011

 ⁶ From January 17, 2012

 ⁷ To December 31, 2011

 ⁸ From September 1, 2011

 ⁹ From July 12, 2011

 ¹⁰ From February 1, 2012

 ¹¹ From April 2, 2012

 ¹² To September 22, 2011

 ¹³ To September 9, 2011

¹⁴ From April 9, 2012
¹⁵ To August 31, 2011
¹⁶ To December 20, 2011
¹⁷ From September 1, 2011
¹⁸ From April 1, 2012
¹⁹ November 2011
²⁰ Joint appointment with Geophysical Laboratory
²¹ From September 2, 2011
²² To May 2, 2012
*In residence for a portion

of the reporting year

EMBRYOLOGY

Avraham-Davidi I., Y. Ely, V. N. Pham, D. Castranova, B. Lo, G. Allmog, T. T. Chen, J. Ungos, K. Kidd, K. Shaw, S. A. Farber, G. S. Shelness, M. L. Iruela-Arispe, B. M. Weinstein, and K. Yaniv, Lipoprotein levels regulate angiogenesis by modulating expression of VEGFR1, *Nat. Med. 18*, 967-973, 2012.

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GEOPHYSICAL LABORATORY

Here updated through September 1, 2012. The list is regularly updated on the Geophysical Laboratory web site (http://www.gl.ciw.edu).

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